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# **CALCEPH - Python language**

***Release 3.5.2***

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This manual documents how to install and use the CALCEPH Library using the Python interface.

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## INTRODUCTION

The CALCEPH Library is designed to access the binary planetary ephemeris files, such INPOPxx and JPL DExxx ephemeris files, (called 'original JPL binary' or 'INPOP 2.0 or 3.0 binary' ephemeris files in the next sections) and the SPICE kernel files (called 'SPICE' ephemeris files in the next sections). At the moment, supported SPICE files are :

- text Planetary Constants Kernel (KPL/PCK) files
- binary PCK (DAF/PCK) files.
- binary SPK (DAF/SPK) files containing segments of type 1, 2, 3, 5, 8, 9, 12, 13, 17, 18, 19, 20, 21, 102, 103 and 120.
- meta kernel (KPL/MK) files.
- frame kernel (KPL/FK) files. Only a basic support is provided.

This library provides a C interface and, optionally, the Fortran 77 or 2003, Python and Octave/Matlab interfaces, to be called by the application.

This library could access to the following ephemeris

- INPOP06 or later
- DE200
- DE403 or later
- EPM2011 or later

Although computers have different endianness (order in which integers are stored as bytes in computer memory), the library could handle the binary ephemeris files with any endianness. This library automatically swaps the bytes when it performs read operations on the ephemeris file.

The internal format of the original JPL binary planetary ephemeris files is described in the paper :

- David Hoffman : 1998, A Set of C Utility Programs for Processing JPL Ephemeris Data, <ftp://ssd.jpl.nasa.gov/pub/eph/export/C-versions/hoffman/EphemUtilVer0.1.tar>

The 'INPOP 2.0 binary' file format for planetary ephemeris files is described in the paper :

- M. Gastineau, J. Laskar, A. Fienga, H. Manche : 2012, INPOP binary ephemeris file format - version 2.0 [http://www.imcce.fr/inpop/inpop\\_file\\_format\\_2\\_0.pdf](http://www.imcce.fr/inpop/inpop_file_format_2_0.pdf)

The 'INPOP 3.0 binary' file format for planetary ephemeris files is described in the paper :

- M. Gastineau, J. Laskar, A. Fienga, H. Manche : 2017, INPOP binary ephemeris file format - version 3.0 [http://www.imcce.fr/inpop/inpop\\_file\\_format\\_3\\_0.pdf](http://www.imcce.fr/inpop/inpop_file_format_3_0.pdf)



## INSTALLATION

The following section describes the installation of the **Python** interface of the library. If you want to install the interface for another programming language, you have to follow the instructions of the manual of that language.

### 2.1 Instructions

The python interface of the library should be installed using the package management system **pip** on all operating systems : Windows and Unix-like system (Linux, Mac OS X, BSD, Cygwin, ...).

A python interpreter, compliant at least with the Python 2.6 or Python 3.0 specifications, and the package Cython, setuptools and numpy are required to install the python interface of the library.

Some Linux distributions require the installation development tools of the python software. The name of this package may change on other Linux distributions or operating systems :

- Debian or Ubuntu distributions : package **python3-dev** and **python-dev-is-python3** are required.
- Opensuse distribution : package **python3-dev** is required.
- Fedora or RedHat distributions : **python3-devel** is required.

In addition you need a C compiler :

- On Unix-like operating systems (Linux, MacOS X, ...), you should install gcc or clang.
- On windows operating system, you need the Microsoft Visual C++ compiler. If you don't have a C compiler already installed, you can download the community edition of visual studio <https://visualstudio.microsoft.com/fr/vs/features/cplusplus/>. Before the execution of the next steps, you should execute the following line in the same terminal. You may have to adjust the path according to your version of the Visual studio compiler.

```
"C:\Program Files (x86)\Microsoft Visual Studio\2019\Community\VC\Auxiliary\Build\  
↪vcvars64.bat"
```

### 2.1.1 Using pip

Depending on your python installation, the command **pip** may be replaced by **pip3**. If you use the distribution *anaconda*, you should prefer to use the instruction from the Anaconda section.

The steps are :

- Install the requirements

```
pip install Cython setuptools numpy
```

- Install the library

```
pip install calcephpy
```

### 2.1.2 Using Anaconda

Depending on your anaconda installation, the command **pip** may be replaced by **pip3**.

The steps are :

- Install the gcc compiler from the Anaconda compiler tools :

See the instruction on <https://docs.conda.io/projects/conda-build/en/latest/resources/compiler-tools.html>

For example, on the operating system Linux, it will be :

```
conda install gcc_linux-64
```

- Install the other requirements

```
conda install Cython setuptools numpy
```

- Install the library

```
pip install calcephpy
```

## LIBRARY INTERFACE

### 3.1 A simple example program

The following example program shows the typical usage of the Python interface.

Other examples using the Python interface can be found in the directory *examples* of the library sources.

```
from calcephpy import *
peph = CalcephBin.open("example1.dat")
AU = peph.getconstant("AU")
jd0 = 2451542
dt = 0.5
PV = peph.compute_unit(jd0, dt, NaifId.MOON, NaifId.EARTH,
                       Constants.UNIT_KM+Constants.UNIT_SEC+Constants.USE_NAIFID)
print(PV)
peph.close()
```

### 3.2 Modules

It is designed to work with Python interpreters compliant with the Python 2.6 or later and Python 3.0 or later.

All declarations needed to use CALCEPH Library are collected in the module `calcephpy`. You should import this module:

```
from calcephpy import *
```

If you receive the following message `ImportError: No module named calcephpy` and if the configuration option `enable-python-package-system` and `enable-python-package-user` was not set, the environment variable `PYTHONPATH` should be set to the right location of the CALCEPH python package ( e.g., `PYTHONPATH=/usr/local/lib64/python3.4/site-packages/:$PYTHONPATH` ) in your shell initialization file (e.g., `~/.bash_login` or `~/.profile`), in order that the python interpreter finds the CALCEPH python package.

Relative to C or Fortran interface, the prefixes `calceph_`, `CALCEPH_`, `NAIFID_` are deleted for the naming convention of the functions, constants and NAIF identification numbers.

## 3.3 Types

`calcephpy.CalcephBin`

This type contains all information to access an ephemeris file.

`calcephpy.NaifId`

This type contains the NAIF identification numbers.

`calcephpy.Constants`

This type contains all constants defined in the library, except the NAIF identification numbers.

## 3.4 Constants

The following constants are defined in the class **Constants** (or *calcephpy.Constants*).

**VERSION\_MAJOR**

This integer constant defines the major revision of this library. It can be used to distinguish different releases of this library.

**VERSION\_MINOR**

This integer constant defines the minor revision of this library. It can be used to distinguish different releases of this library.

**VERSION\_PATCH**

This integer constant defines the patch level revision of this library. It can be used to distinguish different releases of this library.

**VERSION\_STRING**

This string is the version of the library, which can be compared to the result of `calceph_getversion` to check at run time if the header file and library used match:

Note: Obtaining different strings is not necessarily an error, as in general, a program compiled with some old CALCEPH version can be dynamically linked with a newer CALCEPH library version (if allowed by the operating system).

**ASTEROID**

This integer defines the offset value for the asteroids that must be used as target or center for the computation functions, such as `calcephpy.CalcephBin.compute()`.

The following constants specify in which units are expressed the output of the computation functions, such as `calcephpy.CalcephBin.compute_unit()`:

**UNIT\_AU**

This integer defines that the unit of the positions and velocities is expressed in astronomical unit.

**UNIT\_KM**

This integer defines that the unit of the positions and velocities is expressed in kilometer.

#### **UNIT\_DAY**

This integer defines that the unit of the velocities or the quantity TT-TDB or TCG-TCB is expressed in day (one day=86400 seconds).

#### **UNIT\_SEC**

This integer defines that the unit of the velocities or the quantity TT-TDB or TCG-TCB is expressed in second.

#### **UNIT\_RAD**

This integer defines that the unit of the angles is expressed in radian.

#### **OUTPUT\_EULERANGLES**

This integer defines that the output array contains the euler angles.

#### **OUTPUT\_NUTATIONANGLES**

This integer defines that the output array contains the nutation angles.

#### **USE\_NAIFID**

This integer defines that the NAIF identification numbers are used as target or center for the computation functions, such as `calcephpy.CalcephBin.compute_unit()`.

The following constants specify the type of segments for the functions, such as `calcephpy.getmaxsupportedorder()` :

#### **SEGTYPE\_ORIG\_0**

This integer defines the type of segment for the original INPOP/JPL DE file format.

#### **SEGTYPE\_SPK\_1**

#### **SEGTYPE\_SPK\_2**

#### **SEGTYPE\_SPK\_3**

#### **SEGTYPE\_SPK\_5**

#### **SEGTYPE\_SPK\_8**

#### **SEGTYPE\_SPK\_9**

#### **SEGTYPE\_SPK\_12**

#### **SEGTYPE\_SPK\_13**

#### **SEGTYPE\_SPK\_17**

#### **SEGTYPE\_SPK\_18**

#### **SEGTYPE\_SPK\_19**

#### **SEGTYPE\_SPK\_20**

#### **SEGTYPE\_SPK\_21**

#### **SEGTYPE\_SPK\_102**

#### **SEGTYPE\_SPK\_103**

**SEGTYPE\_SPK\_120**

This integer defines the type of segments (1, 2, 3, 5, 8, 9, 12, 13, 17, 18, 19, 20, 21, 102, 103 and 120) for the SPICE Kernel files.

## MULTIPLE FILE ACCESS FUNCTIONS

The following group of functions should be the preferred method to access to the library. They allow to access to multiple ephemeris files at the same time, even by multiple threads.

When an error occurs, these functions execute error handlers according to the behavior defined by the function *calcephpy.seterrorhandler()*.

### 4.1 Time notes

The functions *calcephpy.CalcephBin.compute()*, *calcephpy.CalcephBin.compute\_unit()*, *calcephpy.CalcephBin.compute\_order()*, *calcephpy.CalcephBin.orient\_unit()*, ... only accept a date expressed in the same timescale as the ephemeris files, which can be retrieved using the function *calcephpy.CalcephBin.gettimescale()*. Ephemeris files are generally expressed using the timescale TDB. If a date, expressed in the TT (Terrestrial Time) timescale, is supplied to them, these functions will return an erroneous position of the order of several tens of meters for the planets. If a date, expressed in the Coordinated Universal Time (UTC), is supplied to them, these functions will return a very large erroneous position over several thousand kilometers for the planets.

### 4.2 Thread notes

If the standard I/O functions such as **fread** are not reentrant then the CALCEPH I/O functions using them will not be reentrant either.

It's not safe for two threads to call the functions with the same object of type *CalcephBin* if and only if the function *calcephpy.CalcephBin.isthreadsafe()* returns a non-zero value. A previous call to the function *calcephpy.CalcephBin.prefetch()* is required for the function *calcephpy.CalcephBin.isthreadsafe()* to return a non-zero value.

It's safe for two threads to access simultaneously to the same ephemeris file with two different objects of type *CalcephBin*. In this case, each thread must open the same file.

## 4.3 Usage

The following examples, that can be found in the directory *examples* of the library sources, show the typical usage of this group of functions.

The example in Python language is `pymultiple.py`.

## 4.4 Functions

### 4.4.1 `calcephpy.CalcephBin.open`

`calcephpy.CalcephBin.open(filename) → eph`

**Parameters**

**filename** (*str*) -- pathname of the file

**Returns**

ephemeris descriptor

**Return type**

`calcephpy.CalcephBin`

This function opens the file whose pathname is the string pointed to by `filename`, reads the two header blocks of this file and returns an ephemeris descriptor associated to it. This file must be compliant to the format specified by the 'original JPL binary', 'INPOP 2.0 binary' or 'SPICE' ephemeris file. At the moment, supported SPICE files are the following :

- text Planetary Constants Kernel (KPL/PCK) files
- binary PCK (DAF/PCK) files.
- binary SPK (DAF/SPK) files containing segments of type 1, 2, 3, 5, 8, 9, 12, 13, 17, 18, 19, 20, 21, 102, 103 and 120.
- meta kernel (KPL/MK) files.
- frame kernel (KPL/FK) files. Only a basic support is provided.

Just after the call of `calcephpy.CalcephBin.open()`, the function `calcephpy.CalcephBin.prefetch()` should be called to accelerate future computations.

The function `calcephpy.CalcephBin.close()` must be called to free allocated memory by this function.

The following example opens the ephemeris file `example1.dat`

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")

# ... computation ...

peph.close()
```

### 4.4.2 calcephpy.CalcephBin.open

calcephpy.CalcephBin.open(array\_filename) → eph

**Parameters**

**array\_filename** (*list*) -- array of pathname of the files

**Returns**

ephemeris descriptor

**Return type**

calcephpy.CalcephBin

This function opens *n* files whose pathnames are the string pointed to by *array\_filename*, reads the header blocks of these files and returns an ephemeris descriptor associated to them.

These files must have the same type (e.g., all files are SPICE files or original JPL files). This file must be compliant to the format specified by the 'original JPL binary', 'INPOP 2.0 or 3.0 binary' or 'SPICE' ephemeris file. At the moment, supported SPICE files are the following :

- text Planetary Constants Kernel (KPL/PCK) files
- binary PCK (DAF/PCK) files.
- binary SPK (DAF/SPK) files containing segments of type 1, 2, 3, 5, 8, 9, 12, 13, 17, 18, 19, 20, 21, 102, 103 and 120.
- meta kernel (KPL/MK) files.
- frame kernel (KPL/FK) files. Only a basic support is provided.

Just after the call of `calcephpy.CalcephBin.open()`, the function `calcephpy.CalcephBin.prefetch()` should be called to accelerate future computations.

The function `calcephpy.CalcephBin.close()` must be called to free allocated memory by this function.

The following example opens the ephemeris file `example1.bsp` and `example1.tpc`

```
from calcephpy import CalcephBin
peph = CalcephBin.open(['example1.bsp', 'example1.tpc'])

# ... computation ...

peph.close()
```

### 4.4.3 calcephpy.CalcephBin.prefetch

calcephpy.CalcephBin.prefetch()

This function prefetches to the main memory all files associated to the ephemeris descriptor. This prefetching operation will accelerate the further computations performed with `calcephpy.CalcephBin.compute()`, `calcephpy.CalcephBin.compute_unit()`, `calcephpy.CalcephBin.compute_order()`, `calcephpy.CalcephBin.orient_unit()`, ... .

It requires that the file is smaller than the main memory. If multiple threads (e.g. threads of openMP or Posix Pthreads) prefetch the data for the same ephemeris file, the used memory will remain the same as if the prefetch operation was done by a single thread if and if the endianness of the file is the same as the computer and if the operating system, such as Linux, MacOS X other unix, supports the function `mmap`.

#### 4.4.4 calcephy.CalcephBin.isthreadsafe

calcephy.CalcephBin.isthreadsafe()

This function returns 1 if multiple threads can access the same ephemeris descriptor, otherwise 0.

A previous call to the function `calcephy.CalcephBin.prefetch()` is required, and the library should be compiled with `--enable-thread=yes` on Unix-like operating system, for the function `calcephy.CalcephBin.isthreadsafe()` to return a non-zero value. If the file is not encoded with the same endian as the current hardware, then function may return 0.

If this function returns 1, several threads may use the same ephemeris descriptor for the computational functions `calcephy.CalcephBin.compute()`, .... It allows to use the same object for parallel loops.

#### 4.4.5 calcephy.CalcephBin.compute

calcephy.CalcephBin.compute(*JD0, time, target, center*) → PV

##### Parameters

- **JD0** (*float/list/numpy.ndarray*) -- Integer part of the Julian date (TDB or TCB)
- **time** (*float/list/numpy.ndarray*) -- Fraction part of the Julian date (TDB or TCB)
- **target** (*int*) -- The body or reference point whose coordinates are required (see the list, below).
- **center** (*int*) -- The origin of the coordinate system (see the list, below). If *target* is 14, 15, 16 or 17 (nutaton, libration, TT-TDB or TCG-TCB), *center* must be 0.

##### Returns

Depending on the target value, an array to receive the cartesian position (x,y,z) and the velocity (xdot, ydot, zdot), or a time scale transformation value, or the angles of the librations of the Moon and their derivatives, or the nutation angles and their derivatives.

##### Return type

list

This function reads, if needed, in the ephemeris file *self* and interpolates a single object, usually the position and velocity of one body (*target*) relative to another (*center*) for the time *JD0+time* and stores the results to *PV*. The ephemeris file *self* must have been previously opened with the function `calcephy.CalcephBin.open()`.

The returned array *PV* has the following properties

- If the target is *TT-TDB*, only the first element of this array will get the result. The time scale transformation TT-TDB is expressed in seconds.
- If the target is *TCG-TCB*, only the first element of this array will get the result. The time scale transformation TCG-TCB is expressed in seconds.
- If the target is *Librations*, the array contains the angles of the librations of the Moon and their derivatives. The angles of the librations of the Moon are expressed in radians and their derivatives are expressed in radians per day.
- If the target is *Nutations*, the array contains the nutation angles and their derivatives. The nutation angles are expressed in radians and their derivatives are expressed in radians per day.
- Otherwise the returned values is the cartesian position (x,y,z), expressed in Astronomical Unit (au), and the velocity (xdot, ydot, zdot), expressed in Astronomical Unit per day (au/day).

If *JD0* and *time* are list or NumPy's array (1D) of double-precision floating-point values, the returned array *PV* is a list of 6 arrays. Each array contain a single component of position or velocity (e.g., *PV*[0] contains the coordinate X, *PV*[1] contains the coordinate Y, ...).

The date (*JD0*, *time*) should be expressed in the same timescale as the ephemeris files, which can be retrieved using the function `calcephpy.CalcephBin.gettimescale()`. To get the best numerical precision for the interpolation, the time is splitted in two floating-point numbers. The argument *JD0* should be an integer and *time* should be a fraction of the day. But you may call this function with *time*=0 and *JD0*, the desired time, if you don't take care about numerical precision.

**Warning:** If a date, expressed in the Coordinated Universal Time (UTC), is supplied to this function, a very large erroneous position will be returned.

The possible values for *target* and *center* are :

value	meaning
1	Mercury Barycenter
2	Venus Barycenter
3	Earth
4	Mars Barycenter
5	Jupiter Barycenter
6	Saturn Barycenter
7	Uranus Barycenter
8	Neptune Barycenter
9	Pluto Barycenter
10	Moon
11	Sun
12	Solar Sytem barycenter
13	Earth-moon barycenter
14	Nutation angles
15	Librations
16	TT-TDB
17	TCG-TCB
asteroid number + CALCEPH_asteroid	asteroid

These accepted values by this function are the same as the value for the JPL function *PLEPH*, except for the values *TT-TDB*, *TCG-TCB* and asteroids.

For example, the value "CALCEPH\_asteroid+4" for *target* or *center* specifies the asteroid Vesta.

The following example prints the heliocentric coordinates of Mars at time=2442457.5 and at 2442457.9

```
from calcephpy import *

def printcoord(PV,name):
    print('{0} :\n{1}\n'.format(name,PV))

jd0=2442457
dt1=0.5E0
dt2=0.9E0

peph = CalcephBin.open("example1.dat")
```

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```
PV1 = peph.compute(jd0, dt1, 4, 11)
printcoord(PV1, "heliocentric coordinates of Mars")

PV2 = peph.compute(jd0, dt2, 4, 11)
printcoord(PV2, "heliocentric coordinates of Mars")

peph.close()
```

#### 4.4.6 calcephpy.CalcephBin.compute\_unit

calcephpy.CalcephBin.**compute\_unit**(*JD0, time, target, center, unit*) → PV

##### Parameters

- **JD0** (*float/list/numpy.ndarray*) -- Integer part of the Julian date (TDB or TCB)
- **time** (*float/list/numpy.ndarray*) -- Fraction part of the Julian date (TDB or TCB)
- **target** (*int*) -- The body or reference point whose coordinates are required. The numbering system depends on the parameter unit.
- **center** (*int*) -- The origin of the coordinate system. The numbering system depends on the parameter unit.
- **unit** (*int*) --  
The units of PV.  
This integer is a sum of some unit constants (CALCEPH\_UNIT\_???) and/or the constant *USE\_NAIFID*.  
If the unit contains *USE\_NAIFID*, the NAIF identification numbering system is used for the target and the center (*NAIF identification numbers* for the list).  
If the unit doesnot contain *USE\_NAIFID*, the old number system is used for the target and the center (see the list in the function *calcephpy.CalcephBin.compute()*).

##### Returns

Depending on the target value, an array to receive the cartesian position (x,y,z) and the velocity (xdot, ydot, zdot), or a time scale transformation value, or the angles of the librations of the Moon and their derivatives, or the nutation angles and their derivatives.

##### Return type

list

This function is similar to the function *calcephpy.CalcephBin.compute()*, except that the units of the output are specified.

This function reads, if needed, in the ephemeris file *self* and interpolates a single object, usually the position and velocity of one body (*target*) relative to another (*center*) for the time *JD0+time* and stores the results to *PV*. The ephemeris file *self* must have been previously opened with the function *calcephpy.CalcephBin.open()*. The output values are expressed in the units specified by *unit*.

This function checks the units if invalid combinations of units are given to the function.

The date (JD0, time) should be expressed in the same timescale as the ephemeris files, which can be retrieved using the function *calcephpy.CalcephBin.gettimescale()*.

**Warning:** If a date, expressed in the Coordinated Universal Time (UTC), is supplied to this function, a very large erroneous position will be returned.

The returned array *PV* has the following properties

- If the target is the time scale transformation TT-TDB, only the first element of this array will get the result.
- If the target is the time scale transformation TCG-TCB, only the first element of this array will get the result.
- If the target is *Librations*, the array contains the angles of the librations of the Moon and their derivatives.
- If the target is *Nutations*, the array contains the nutation angles and their derivatives.
- Otherwise the returned value is the cartesian position (x,y,z) and the velocity (xdot, ydot, zdot).

If *JD0* and *time* are list or NumPy's array (1D) of double-precision floating-point values, the returned array *PV* is a list of 6 arrays. Each array contain a single component of position or velocity (e.g., *PV*[0] contains the coordinate X, *PV*[1] contains the coordinate Y, ...) .

The values stored in the array *PV* are expressed in the following units

- The position and velocity are expressed in Astronomical Unit (au) if unit contains *UNIT\_AU*.
- The position and velocity are expressed in kilometers if unit contains *UNIT\_KM*.
- The velocity, TT-TDB, TCG-TCB, the derivatives of the angles of the nutation, or the derivatives of the librations of the Moon or are expressed in days if unit contains *UNIT\_DAY*.
- The velocity, TT-TDB, TCG-TCB, the derivatives of the angles of the nutation, or the derivatives of the librations of the Moon are expressed in seconds if unit contains *UNIT\_SEC*.
- The angles of the librations of the Moon or the nutation angles are expressed in radians if unit contains *UNIT\_RAD*.

For example, to get the position and velocities expressed in kilometers and kilometers/seconds, the unit must be set to *UNIT\_KM + UNIT\_SEC*.

The following example prints the heliocentric coordinates of Mars at time=2442457.5

```
from calcephpy import *

def printcoord(PV,name):
    print('{0} :\n{1}\n'.format(name,PV))

jd0=2442457
dt=0.5E0

peph = CalcephBin.open("example1.dat")

PV1 = peph.compute_unit(jd0, dt, 4, 11, Constants.UNIT_KM+Constants.UNIT_SEC)
printcoord(PV1,"heliocentric coordinates of Mars")

PV2 = peph.compute_unit(jd0, dt, NaifId.MARS_BARYCENTER, NaifId.SUN,
                        Constants.UNIT_KM+Constants.UNIT_SEC+Constants.USE_NAIFID)
printcoord(PV2,"heliocentric coordinates of Mars")

peph.close()
```

#### 4.4.7 calcephpy.CalcephBin.orient\_unit

calcephpy.CalcephBin.orient\_unit(JD0, time, target, unit) → PV

##### Parameters

- **JD0** (*float/list/numpy.ndarray*) -- Integer part of the Julian date (TDB or TCB)
- **time** (*float/list/numpy.ndarray*) -- Fraction part of the Julian date (TDB or TCB)
- **target** (*int*) -- The body whose orientations are requested. The numbering system depends on the parameter unit.
- **unit** (*int*) --

The units of PV.

This integer is a sum of some unit constants (CALCEPH\_UNIT\_???) and/or the constant [USE\\_NAIFID](#).

If the unit contains [USE\\_NAIFID](#), the NAIF identification numbering system is used for the target (*NAIF identification numbers* for the list).

If the unit does not contain [USE\\_NAIFID](#), the old number system is used for the target (see the list in the function [calcephpy.CalcephBin.compute\(\)](#)).

##### Returns

An array to receive the euler angles, or nutation angles, and their derivatives for the orientation of the body.

##### Return type

list

This function reads, if needed, in the ephemeris file *self* and interpolates the orientation of a single body (*target*) for the time *JD0+time* and stores the results to *PV*. The ephemeris file *self* must have been previously opened with the function [calcephpy.CalcephBin.open\(\)](#). The output values are expressed in the units specified by *unit*.

The date (JD0, time) should be expressed in the same timescale as the ephemeris files, which can be retrieved using the function [calcephpy.CalcephBin.gettimescale\(\)](#).

This function checks the units if invalid combinations of units are given to the function.

The returned array *PV* has the following properties

- If *unit* contains [OUTPUT\\_NUTATIONANGLES](#), the array contains the nutation angles and their derivatives for the orientation of the body. At the present moment, only the nutation for the earth are supported in the original DE files.
- If *unit* contains [OUTPUT\\_EULERANGLES](#), or doesnot contain [OUTPUT\\_NUTATIONANGLES](#), the array contains the euler angles and their derivatives for the orientation of the body.

If *JD0* and *time* are list or NumPy's array (1D) of double-precision floating-point values, the returned array *PV* is a list of 6 arrays. Each array contain a single component of orientation.

The values stored in the array *PV* are expressed in the following units

- The derivatives of the angles are expressed in days if unit contains [UNIT\\_DAY](#).
- The derivatives of the angles are expressed in seconds if unit contains [UNIT\\_SEC](#).
- The angles and their derivatives are expressed in radians if unit contains [UNIT\\_RAD](#).

For example, to get the nutation angles of the Earth and their derivatives expressed in radian and radian/seconds using the NAIF identification numbering system, the target must be set to NAIFID\_EARTH and the unit must be set to [OUTPUT\\_NUTATIONANGLES](#) + [UNIT\\_RAD](#) + [UNIT\\_SEC](#).

The following example prints the angles of libration of the Moon at time=2442457.5

```

from calcephpy import *

jd0=2442457
dt=0.5E0

peph = CalcephBin.open("example1.dat")

PV = peph.orient_unit(jd0, dt, NaifId.MOON,
                      Constants.USE_NAIFID+Constants.UNIT_RAD+Constants.UNIT_SEC)
print(PV)

peph.close()

```

#### 4.4.8 calcephpy.CalcephBin.rotangmom\_unit

calcephpy.CalcephBin.**rotangmom\_unit**(JD0, time, target, unit) → PV

##### Parameters

- **JD0** (*float*) -- Integer part of the Julian date (TDB or TCB)
- **time** (*float*) -- Fraction part of the Julian date (TDB or TCB)
- **target** (*int*) -- The body whose orientations are requested. The numbering system depends on the parameter unit.
- **unit** (*int*) --

The units of PV.

This integer is a sum of some unit constants (CALCEPH\_UNIT\_???) and/or the constant [USE\\_NAIFID](#).

If the unit contains [USE\\_NAIFID](#), the NAIF identification numbering system is used for the target (*NAIF identification numbers* for the list).

If the unit does not contain [USE\\_NAIFID](#), the old number system is used for the target (see the list in the function [calcephpy.CalcephBin.compute\(\)](#)).

##### Returns

An array to receive the angular momentum due to its rotation, divided by the product of the mass and of the square of the radius, and the derivatives, of the body.

##### Return type

list

This function reads, if needed, in the ephemeris file *self* and interpolates the angular momentum vector due to the rotation of the body, divided by the product of the mass  $m$  and of the square of the radius  $R$ , of a single body (*target*) for the time  $JD0+time$  and stores the results to *PV*. The ephemeris file *self* must have been previously opened with the function [calcephpy.CalcephBin.open\(\)](#). The angular momentum  $L$ , due to the rotation of the body, is defined as the product of the inertia matrix  $I$  by the angular velocity vector  $\omega$ . So the returned value is  $L/(mR^2) = (I\omega)/(mR^2)$

The date (JD0, time) should be expressed in the same timescale as the ephemeris files, which can be retrieved using the function [calcephpy.CalcephBin.gettimescale\(\)](#).

The output values are expressed in the units specified by *unit*.

This function checks the units if invalid combinations of units are given to the function.

The values stored in the array *PV* are expressed in the following units

- The angular momentum and its derivative are expressed in days if unit contains [UNIT\\_DAY](#).
- The angular momentum and its derivative are expressed in seconds if unit contains [UNIT\\_SEC](#).

The following example prints the angular momentum, due to its rotation, for the Earth at time=2451419.5

```
from calcephpy import *

jd0=2451419
dt=0.5E0

peph = CalcephBin.open("example2_rotangmom.dat")

G = peph.rotangmom_unit(jd0, dt, NaifId.EARTH,
                        Constants.USE_NAIFID+Constants.UNIT_SEC)
print(G)

peph.close()
```

#### 4.4.9 calcephpy.CalcephBin.compute\_order

calcephpy.CalcephBin.**compute\_order**(*JD0, time, target, center, unit, order*) → PVAJ

##### Parameters

- **JD0** (*float/list/numpy.ndarray*) -- Integer part of the Julian date (TDB or TCB)
- **time** (*float/list/numpy.ndarray*) -- Fraction part of the Julian date (TDB or TCB)
- **target** (*int*) -- The body or reference point whose coordinates are required. The numbering system depends on the parameter unit.
- **center** (*int*) -- The origin of the coordinate system. The numbering system depends on the parameter unit.
- **unit** (*int*) --  
The units of PVAJ.  
This integer is a sum of some unit constants (CALCEPH\_UNIT\_???) and/or the constant [USE\\_NAIFID](#).  
If the unit contains [USE\\_NAIFID](#), the NAIF identification numbering system is used for the target and the center (*NAIF identification numbers* for the list).  
If the unit doesnot contain [USE\\_NAIFID](#), the old number system is used for the target and the center (see the list in the function [calcephpy.CalcephBin.compute\(\)](#)).
- **order** (*int*) -- The order of derivatives
  - = 0 , only the position is computed. The first three numbers of PVAJ are valid for the results.
  - = 1 , only the position and velocity are computed. The first six numbers of PVAJ are valid for the results.
  - = 2 , only the position, velocity and acceleration are computed. The first nine numbers of PVAJ are valid for the results.
  - = 3 , the position, velocity and acceleration and jerk are computed. The first twelve numbers of PVAJ are valid for the results.

If order equals to 1, the behavior of `calcephpy.CalcephBin.compute_order()` is the same as `calcephpy.CalcephBin.compute_unit()`.

### Returns

Depending on the target value, an array to receive the cartesian position (x,y,z), the velocity (xdot, ydot, zdot), the acceleration and the jerk, or a time scale transformation value, or the angles of the librations of the Moon and their successive derivatives, or the nutation angles and their successive derivatives.

### Return type

list

This function is similar to the function `calcephpy.CalcephBin.compute_unit()`, except that the order of the computed derivatives is specified.

This function reads, if needed, in the ephemeris file *self* and interpolates a single object, usually the position and their derivatives of one body (*target*) relative to another (*center*) for the time *JD0+time* and stores the results to *PVAJ*. The ephemeris file *self* must have been previously opened with the function `calcephpy.CalcephBin.open()`. The order of the derivatives are specified by *order*. The output values are expressed in the units specified by *unit*.

The returned array *PVAJ* has the following properties

- If the target is the time scale transformation TT-TDB, only the first elements of each component will get the result.
- If the target is the time scale transformation TCG-TCB, only the first elements of each component will get the result.
- If the target is *Librations*, the array contains the angles of the librations of the Moon and their successive derivatives.
- If the target is *Nutations*, the array contains the nutation angles and their successive derivatives.
- Otherwise the returned value is the cartesian position (x,y,z), the velocity (xdot, ydot, zdot), the jerk and the acceleration.

The returned array *PVAJ* must be large enough to store the results.

- *PVAJ*[1:3] contain the position (x,y,z) and is always valid.
- *PVAJ*[4:6] contain the velocity (dx/dt,dy/dt,dz/dt) and is only valid if *order* is greater or equal to 1.
- *PVAJ*[7:9] contain the acceleration ( $d^2x/dt^2, d^2y/dt^2, d^2z/dt^2$ ) and is only valid if *order* is greater or equal to 2.
- *PVAJ*[10:12] contain the jerk ( $d^3x/dt^3, d^3y/dt^3, d^3z/dt^3$ ) and is only valid if *order* is equal to 3.

If *JD0* and *time* are list or NumPy's array (1D) of double-precision floating-point values, the returned array *PVAJ* is a list of  $3*(order+1)$  arrays. Each array contain a single component of position, velocity ... (e.g., *PV*[0] contains the coordinate X, *PV*[1] contains the coordinate Y, ...) .

The date (*JD0*, *time*) should be expressed in the same timescale as the ephemeris files, which can be retrieved using the function `calcephpy.CalcephBin.gettimescale()`.

**Warning:** If a date, expressed in the Coordinated Universal Time (UTC), is supplied to this function, a very large erroneous position will be returned.

The values stored in the array *PVAJ* are expressed in the following units

- The position, velocity, acceleration and jerk are expressed in Astronomical Unit (au) if unit contains `UNIT_AU`.
- The position, velocity, acceleration and jerk are expressed in kilometers if unit contains `UNIT_KM`.

- The velocity, acceleration, jerk, TT-TDB, TCG-TCB or the derivatives of the angles of the librations of the Moon are expressed in days if unit contains `UNIT_DAY`.
- The velocity, acceleration, jerk, TT-TDB, TCG-TCB or the derivatives of the angles of the librations of the Moon are expressed in seconds if unit contains `UNIT_SEC`.
- The angles of the librations of the Moon are expressed in radians if unit contains `UNIT_RAD`.

For example, to get the positions, velocities, accelerations and jerks expressed in kilometers and kilometers/seconds, the unit must be set to `UNIT_KM + UNIT_SEC`.

This function checks the units if invalid combinations of units are given to the function.

The following example prints the heliocentric coordinates of Mars at time=2442457.5

```
from calcephpy import *

jd0=2442457
dt=0.5E0

peph = CalcephBin.open("example1.dat")

# compute only the heliocentric position of Mars in km
P = peph.compute_order(jd0, dt, NaifId.MARS_BARYCENTER, NaifId.SUN,
                      Constants.UNIT_KM+Constants.UNIT_SEC+Constants.USE_NAIFID, 0)
print(P)

# compute positions, velocities, accelerations and jerks of Mars in km and seconds
PVAJ = peph.compute_order(jd0, dt, NaifId.MARS_BARYCENTER, NaifId.SUN,
                          Constants.UNIT_KM+Constants.UNIT_SEC+Constants.USE_NAIFID, 3)
print(PVAJ)

peph.close()
```

#### 4.4.10 calcephpy.CalcephBin.orient\_order

`calcephpy.CalcephBin.orient_order(JD0, time, target, unit, order) → PVAJ`

##### Parameters

- **JD0** (*float/list/numpy.ndarray*) -- Integer part of the Julian date (TDB or TCB)
- **time** (*float/list/numpy.ndarray*) -- Fraction part of the Julian date (TDB or TCB)
- **target** (*int*) -- The body whose orientations are requested. The numbering system depends on the parameter unit.
- **unit** (*int*) --  
The units of PV.  
This integer is a sum of some unit constants (`CALCEPH_UNIT_???`) and/or the constant `USE_NAIFID`.  
If the unit contains `USE_NAIFID`, the NAIF identification numbering system is used for the target (*NAIF identification numbers* for the list).  
If the unit does not contain `USE_NAIFID`, the old number system is used for the target (see the list in the function `calcephpy.CalcephBin.compute()`).
- **order** (*int*) -- The order of derivatives.

- = 0 , only the angles is computed. The first three numbers of PVAJ are valid for the results.
- = 1 , only the angles and the first derivative are computed. The first six numbers of PVAJ are valid for the results.
- = 2 , only the angles and the first and second derivatives are computed. The first nine numbers of PVAJ are valid for the results.
- = 3 , the angles and the first, second and third derivatives are computed. The first twelve numbers of PVAJ are valid for the results.

If order equals to 1, the behavior of `calcephy.CalcephBin.orient_order()` is the same as `calcephy.CalcephBin.orient_unit()`.

#### Returns

An array to receive the euler angles, or nutation angles, and their derivatives for the orientation of the body.

#### Return type

list

This function is similar to the function `calcephy.CalcephBin.orient_unit()`, except that the order of the computed derivatives is specified.

This function reads, if needed, in the ephemeris file *self* and interpolates the orientation of a single body (*target*) for the time *JD0+time* and stores the results to *PVAJ*. The order of the derivatives are specified by *order*. The ephemeris file *self* must have been previously opened with the function `calcephy.CalcephBin.open()`. The output values are expressed in the units specified by *unit*.

This function checks the units if invalid combinations of units are given to the function.

The returned array *PVAJ* has the following properties

- If *unit* contains `OUTPUT_NUTATIONANGLES`, the array contains the nutation angles and their successive derivatives for the orientation of the body. At the present moment, only the nutation for the earth are supported in the original DE files.
- If *unit* contains `OUTPUT_EULERANGLES`, or doesnot contain `OUTPUT_NUTATIONANGLES`, the array contains the euler angles and their successive derivatives for the orientation of the body.

The returned array *PVAJ* must be large enough to store the results.

- *PVAJ*[1:3] contain the angles and is always valid.
- *PVAJ*[4:6] contain the first derivative and is only valid if *order* is greater or equal to 1.
- *PVAJ*[7:9] contain the second derivative and is only valid if *order* is greater or equal to 2.
- *PVAJ*[10:12] contain the third derivative and is only valid if *order* is equal to 3.

If *JD0* and *time* are list or NumPy's array (1D) of double-precision floating-point values, the returned array *PVAJ* is a list of 3\*(order+1) arrays. Each array contain a single component of the orientation.

The values stored in the array *PVAJ* are expressed in the following units

- The derivatives of the angles are expressed in days if unit contains `UNIT_DAY`.
- The derivatives of the angles are expressed in seconds if unit contains `UNIT_SEC`.
- The angles and their derivatives are expressed in radians if unit contains `UNIT_RAD`.

The date (JD0, time) should be expressed in the same timescale as the ephemeris files, which can be retrieved using the function `calcephy.CalcephBin.gettimescale()`.

The following example prints only the angles of libration of the Moon at time=2442457.5

```
from calcephpy import *

jd0=2442457
dt=0.5E0

peph = CalcephBin.open("example1.dat")

P = peph.orient_order(jd0, dt, NaifId.MOON,
                      Constants.USE_NAIFID+Constants.UNIT_RAD+Constants.UNIT_SEC, 0)
print(P)

peph.close()
```

#### 4.4.11 calcephpy.CalcephBin.rotangmom\_order

calcephpy.CalcephBin.**rotangmom\_order**(JD0, time, target, unit, order) → PVAJ

##### Parameters

- **JD0** (*float*) -- Integer part of the Julian date (TDB or TCB)
- **time** (*float*) -- Fraction part of the Julian date (TDB or TCB)
- **target** (*int*) -- The body whose orientations are requested. The numbering system depends on the parameter unit.
- **unit** (*int*) --  
The units of PV.  
This integer is a sum of some unit constants (CALCEPH\_UNIT\_???) and/or the constant [USE\\_NAIFID](#).  
If the unit contains [USE\\_NAIFID](#), the NAIF identification numbering system is used for the target (*NAIF identification numbers* for the list).  
If the unit does not contain [USE\\_NAIFID](#), the old number system is used for the target (see the list in the function [calcephpy.CalcephBin.compute\(\)](#)).
- **order** (*int*) -- The order of derivatives.
  - = 0 , only the angular momentum is computed. The first three numbers of PVAJ are valid for the results.
  - = 1 , only the angular momentum and the first derivative are computed. The first six numbers of PVAJ are valid for the results.
  - = 2 , only the angular momentum and the first and second derivatives are computed. The first nine numbers of PVAJ are valid for the results.
  - = 3 , the angular momentum and the first, second and third derivatives are computed. The first twelve numbers of PVAJ are valid for the results.If order equals to 1, the behavior of [calcephpy.CalcephBin.rotangmom\\_order\(\)](#) is the same as [calcephpy.CalcephBin.rotangmom\\_unit\(\)](#).

##### Returns

An array to receive the angular momentum due to its rotation, divided by the product of the mass and of the square of the radius, and their different order of the derivatives, of the body.

**Return type**

list

This function is similar to the function `calcephpy.CalcephBin.orient_unit()`, except that the order of the computed derivatives is specified.

This function reads, if needed, in the ephemeris file *self* and interpolates the angular momentum vector due to the rotation of the body, divided by the product of the mass  $m$  and of the square of the radius  $R$ , of a single body (*target*) for the time  $JD0+time$  and stores the results to *PVAJ*. The angular momentum  $L$ , due to the rotation of the body, is defined as the product of the inertia matrix  $I$  by the angular velocity vector  $\omega$ . So the returned value is  $L/(mR^2) = (I\omega)/(mR^2)$ . The order of the derivatives are specified by *order*. The ephemeris file *self* must have been previously opened with the function `calcephpy.CalcephBin.open()`. The output values are expressed in the units specified by *unit*.

This function checks the units if invalid combinations of units are given to the function.

The returned array *PVAJ* must be large enough to store the results.

- *PVAJ*[1:3] contain the angular momentum and is always valid.
- *PVAJ*[4:6] contain the first derivative and is only valid if *order* is greater or equal to 1.
- *PVAJ*[7:9] contain the second derivative and is only valid if *order* is greater or equal to 2.
- *PVAJ*[10:12] contain the third derivative and is only valid if *order* is equal to 3.

The values stored in the array *PVAJ* are expressed in the following units

- The angular momentum and its derivatives are expressed in days if unit contains `UNIT_DAY`.
- The angular momentum and its derivatives are expressed in seconds if unit contains `UNIT_SEC`.

The date (JD0, time) should be expressed in the same timescale as the ephemeris files, which can be retrieved using the function `calcephpy.CalcephBin.gettimescale()`.

The following example prints only the angular momentum, due to its rotation, of the Earth at time=2451419.5

```
from calcephpy import *

jd0=2451419
dt=0.5E0

peph = CalcephBin.open("example2_rotangmom.dat")

G = peph.rotangmom_order(jd0, dt, NaifId.EARTH,
                        Constants.USE_NAIFID+Constants.UNIT_SEC, 0)

print(G)

peph.close()
```

#### 4.4.12 `calcephy.CalcephBin.getconstant`

`calcephy.CalcephBin.getconstant(name) → value`

**Parameters**

**name** (*str*) -- name of the constant

**Returns**

first value of the constant

**Return type**

float

This function returns the value associated to the constant *name* in the header of the ephemeris file *self*. Only the first value is returned if multiple values are associated to a constant, such as a list of values.

This function is the same function as `calcephy.CalcephBin.getconstantsd()`.

The following example prints the value of the astronomical unit stored in the ephemeris file

```
from calcephy import *

peph = CalcephBin.open("example1.dat")
AU = peph.getconstant("AU")
print(AU)
peph.close()
```

#### 4.4.13 `calcephy.CalcephBin.getconstantsd`

`calcephy.CalcephBin.getconstantsd(name) → value`

**Parameters**

**name** (*str*) -- name of the constant

**Returns**

first value of the constant

**Return type**

float

This function returns, as a floating-point number, the value associated to the constant *name* in the header of the ephemeris file *self*. Only the first value is returned if multiple values are associated to a constant, such as a list of values. The value must be a floating-point or integer number, otherwise an error is reported.

This function is the same function as `calcephy.CalcephBin.getconstant()`.

The following example prints the value of the astronomical unit stored in the ephemeris file

```
from calcephy import *

peph = CalcephBin.open("example1.dat")
AU = peph.getconstantsd("AU")
print(AU)
peph.close()
```

#### 4.4.14 calcephpy.CalcephBin.getconstantvd

calcephpy.CalcephBin.getconstantvd(*name*) → arrayvalue

**Parameters**

**name** (*str*) -- name of the constant

**Returns**

array of values for the constant

**Return type**

list

This function returns, as floating-point numbers, all values associated to the constant *name* in the header of the ephemeris file *self*.

The values must be floating-point or integer numbers, otherwise an error is reported.

The following example prints the body radii of the earth stored in the ephemeris file

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
radii = peph.getconstantvd("BODY399_RADII")
print(radii)
peph.close()
```

#### 4.4.15 calcephpy.CalcephBin.getconstantss

calcephpy.CalcephBin.getconstantss(*name*) → value

**Parameters**

**name** (*str*) -- name of the constant

**Returns**

first value of the constant

**Return type**

str

This function returns, as a string of character, the value associated to the constant *name* in the header of the ephemeris file *self*. Only the first value is returned if multiple values are associated to a constant, such as a list of values. The value must be a string, otherwise an error is reported.

The following example prints the value of the unit stored in the ephemeris file

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
UNIT = peph.getconstantss("UNIT")
print(UNIT)
peph.close()
```

#### 4.4.16 calcephpy.CalcephBin.getconstantvs

calcephpy.CalcephBin.getconstantvs(*name*) → arrayvalue

**Parameters**

**name** (*str*) -- name of the constant

**Returns**

array of values for the constant

**Return type**

list

This function returns, as strings of characters, all values associated to the constant *name* in the header of the ephemeris file *self*.

The values must be strings, otherwise an error is reported.

The following example prints the units of the mission stored in the ephemeris file

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
mission_units = peph.getconstantvs("MISSION_UNITS")
print(mission_units)
peph.close()
```

#### 4.4.17 calcephpy.CalcephBin.getconstantcount

calcephpy.CalcephBin.getconstantcount()

**Returns**

number of constants

**Return type**

int

This function returns the number of constants available in the header of the ephemeris file *self*.

The following example prints the number of available constants stored in the ephemeris file

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
n = peph.getconstantcount()
print("number of constants", n)
peph.close()
```

#### 4.4.18 calcephpy.CalcephBin.getconstantindex

calcephpy.CalcephBin.getconstantindex(index) → name, value

**Parameters**

**index** (int) -- index of the constant, between 1 and `calcephpy.CalcephBin.getconstantcount()`

**Returns**

name of the constant, first value of the constant

**Return type**

str, float

This function returns the name and its value of the constant available at the specified index in the header of the ephemeris file *self*. The value of *index* must be between 1 and `calcephpy.CalcephBin.getconstantcount()`.

Only the first value is returned if multiple values are associated to a constant, such as a list of values. If the first value is not an floating-point number, such as a string, then the function returns 0 without raising an error.

The following example displays the name of the constants, stored in the ephemeris file, and their values

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
n = peph.getconstantcount()
for j in range(1, n+1):
    name, value = peph.getconstantindex(j)
    print(name, value)

peph.close()
```

#### 4.4.19 calcephpy.CalcephBin.getfileversion

calcephpy.CalcephBin.getfileversion()

**Returns**

version of the ephemeris file

**Return type**

str

This function returns the version of the ephemeris file, as a string. For example, the argument version will contain 'INPOP10B', 'EPM2017' or 'DE405', ... .

If the file is an original JPL binary planetary ephemeris, then the version of the file can always be determined. If the file is a spice kernel, the version of the file is retrieved from the constant *INPOP\_PCK\_VERSION*, *EPM\_PCK\_VERSION*, or *PCK\_VERSION*.

The following example prints the version of the ephemeris file.

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
version = peph.getfileversion()
print(version)
peph.close()
```

#### 4.4.20 calcephpy.CalcephBin.gettimescale

calcephpy.CalcephBin.gettimescale()

**Returns**

time scale of the ephemeris file

**Return type**

int

This function returns the timescale of the ephemeris file *self* :

- 1 if the quantities of all bodies are expressed in the TDB time scale.
- 2 if the quantities of all bodies are expressed in the TCB time scale.

The following example prints the time scale available in the ephemeris file

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
timescale = peph.gettimescale()
print(timescale)
peph.close()
```

#### 4.4.21 calcephpy.CalcephBin.gettimespan

calcephpy.CalcephBin.gettimespan() → firsttime, lasttime, continuous

**Returns**

first and last available time, availability of the quantities of the bodies over the time span

**Return type**

float, float, int

This function returns the first and last time available in the ephemeris file *self*. The Julian date for the first and last time are expressed in the time scale returned by [calcephpy.CalcephBin.gettimescale\(\)](#) .

It returns the following value in the parameter *continuous* :

- 1 if the quantities of all bodies are available for any time between the first and last time.
- 2 if the quantities of some bodies are available on discontinuous time intervals between the first and last time.
- 3 if the quantities of each body are available on a continuous time interval between the first and last time, but not available for any time between the first and last time.

The following example prints the first and last time available in the ephemeris file

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
firsttime, lasttime, continuous = peph.gettimespan()
print(firsttime, lasttime, continuous)
peph.close()
```

#### 4.4.22 calcephpy.CalcephBin.getpositionrecordcount

calcephpy.CalcephBin.getpositionrecordcount()

**Returns**

number of position's records

**Return type**

int

This function returns the number of position's records available in the ephemeris file *self*. Usually, the number of records is equal to the number of bodies in the ephemeris file if the timespan is continuous. If the timespan is discontinuous for the target and center bodies, then each different timespan is counted as a different record. If the ephemeris file contain timescale transformations' records, such as *TT-TDB* or *TCG-TCB*, then these records are included in the returned value.

The following example prints the number of position's records available in the ephemeris file

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
n = peph.getpositionrecordcount()
print("number of position's record", n)
peph.close()
```

#### 4.4.23 calcephpy.CalcephBin.getpositionrecordindex

calcephpy.CalcephBin.getpositionrecordindex(*index*) → target, center, firsttime, lasttime, frame

**Parameters**

**index** (*int*) -- index of the position's record, between 1 and [calcephpy.CalcephBin.getpositionrecordcount\(\)](#)

**Returns**

**target** : the target body  
**center** : the origin body  
**firsttime** : julian date of the first time  
**lasttime** : julian date of the last time  
**frame** : reference frame (see the list, below)

**Return type**

int, int, float, float, int

This function returns the target and origin bodies, the first and last time, and the reference frame available at the specified index for the position's records of the ephemeris file *self*. The NAIF identification numbering system is used for the target and center integers (*NAIF identification numbers* for the list). The Julian date for the first and last time are expressed in the time scale returned by [calcephpy.CalcephBin.gettimescale\(\)](#).

It returns the following value in the parameter *frame* :

value	Name
1	ICRF

The following example displays the position's records stored in the ephemeris file.

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
n = peph.getpositionrecordcount()
for j in range(1, n+1):
    itarget, icenter, firsttime, lasttime, iframe = peph.getpositionrecordindex(j)
    print(itarget, icenter, firsttime, lasttime, iframe)

peph.close()
```

#### 4.4.24 calcephpy.CalcephBin.getpositionrecordindex2

calcephpy.CalcephBin.getpositionrecordindex2(*index*) → target, center, firsttime, lasttime, frame, segid

##### Parameters

**index** (*int*) -- index of the position's record, between 1 and `calcephpy.CalcephBin.getpositionrecordcount()`

##### Returns

**target** : the target body  
**center** : the origin body  
**firsttime** : julian date of the first time  
**lasttime** : julian date of the last time  
**frame** : reference frame (see the list, below)  
**segid** : segment type (see the details, below)

##### Return type

int, int, float, float, int, int

This function returns the target and origin bodies, the first and last time, the reference frame, and the segment type available at the specified index for the position's records of the ephemeris file *self*. The NAIF identification numbering system is used for the target and center integers (*NAIF identification numbers* for the list). The Julian date for the first and last time are expressed in the time scale returned by `calcephpy.CalcephBin.gettimescale()`.

It returns the following value in the parameter *frame* :

value	Name
1	ICRF

It returns in the parameter *segid* one of the predefined constants `Constants.SEGTYPE_...` (*Constants*).

The following example displays the position's records stored in the ephemeris file.

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
n = peph.getpositionrecordcount()
```

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```

for j in range(1, n+1):
    itarget, icenter, firsttime, lasttime, iframe, iseg = peph.getpositionrecordindex2(j)
    print(itarget, icenter, firsttime, lasttime, iframe, iseg)

peph.close()

```

#### 4.4.25 calcephy.CalcephBin.getorientrecordcount

calcephy.CalcephBin.getorientrecordcount()

##### Returns

number of orientation's records

##### Return type

int

This function returns the number of orientation's records available in the ephemeris file *self*. Usually, the number of records is equal to the number of bodies in the ephemeris file if the timespan is continuous. If the timespan is discontinuous for the target body, then each different timespan is counted as a different record.

The following example prints the number of orientation's records available in the ephemeris file

```

from calcephy import *

peph = CalcephBin.open("example1.dat")
n = peph.getorientrecordcount()
print("number of orientation's record", n)
peph.close()

```

#### 4.4.26 calcephy.CalcephBin.getorientrecordindex

calcephy.CalcephBin.getorientrecordindex(*index*) → target, firsttime, lasttime, frame

##### Parameters

**index** (int) -- index of the orientation's record, between 1 and `calcephy.CalcephBin.getorientrecordcount()`

##### Returns

**target** : the target body  
**firsttime** : julian date of the first time  
**lasttime** : julian date of the last time  
**frame** : reference frame (see the list, below)

##### Return type

int, float, float, int

This function returns the target body, the first and last time, and the reference frame available at the specified index for the orientation's records of the ephemeris file *self*. The NAIF identification numbering system is used for the target

body (*NAIF identification numbers* for the list). The Julian date for the first and last time are expressed in the time scale returned by `calcephpy.CalcephBin.gettimescale()`.

It returns the following value in the parameter *frame* :

value	Name
1	ICRF

The following example displays the orientation's records stored in the ephemeris file.

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
n = peph.getorientrecordcount()
for j in range(1, n+1):
    itarget, firsttime, lasttime, iframe = peph.getorientrecordindex(j)
    print(itarget, firsttime, lasttime, iframe)

peph.close()
```

#### 4.4.27 calcephpy.CalcephBin.getorientrecordindex2

`calcephpy.CalcephBin.getorientrecordindex2(index)` → target, firsttime, lasttime, frame, segid

##### Parameters

**index** (*int*) -- index of the orientation's record, between 1 and `calcephpy.CalcephBin.getorientrecordcount()`

##### Returns

**target** : the target body

**firsttime** : julian date of the first time

**lasttime** : julian date of the last time

**frame** : reference frame (see the list, below)

**segid** : segment type (see the details, below)

##### Return type

int, float, float, int, int

This function returns the target body, the first and last time, the reference frame and the segment type available at the specified index for the orientation's records of the ephemeris file *self*. The NAIF identification numbering system is used for the target body (*NAIF identification numbers* for the list). The Julian date for the first and last time are expressed in the time scale returned by `calcephpy.CalcephBin.gettimescale()`.

It returns the following value in the parameter *frame* :

value	Name
1	ICRF

It returns in the parameter *segid* one of the predefined constants `Constants.SEGTYPE_...` (*Constants*).

The following example displays the orientation's records stored in the ephemeris file.

```
from calcephpy import *

peph = CalcephBin.open("example1.dat")
n = peph.getorientrecordcount()
for j in range(1, n+1):
    itarget, firsttime, lasttime, iframe, iseg = peph.getorientrecordindex2(j)
    print(itarget, firsttime, lasttime, iframe, iseg)

peph.close()
```

#### 4.4.28 calcephpy.CalcephBin.close

calcephpy.CalcephBin.close()

This function closes the access associated to the ephemeris descriptor and frees allocated memory for it.



## ERROR FUNCTIONS

The following group of functions defines the behavior of the library when errors occur during the execution.

### 5.1 Usage

The following examples, that can be found in the directory *examples* of the library sources, show the typical usage of this group of functions.

The example in Python language is `pyerror.py`.

The following example shows how to stop the execution on the error.

```
from calcephpy import *

#set the error handler to stop on error
seterrorhandler(2, 0);

# open the ephemeris file
peph = CalcephBin.open("example1.dat")
```

The following example shows how to define a custom error handler function.

```
from calcephpy import *

#-----
# custom error handler
#-----
def myhandler(msg):
    print("The calceph calls the function myhandler");
    print("The message contains {0} characters\n".format(len(msg)))
    print("The error message is :")
    print("-----")
    print(msg)
    print("-----")
    print("The error handler returns")

# set the error handler to use my own callback
seterrorhandler(3, myhandler)

# open the ephemeris file
peph = CalcephBin.open("example1.dat")
```

## 5.2 calcephpy.seterrorhandler

`calcephpy.seterrorhandler(typehandler, userfunc)`

### Parameters

- **typehandler** (*int*) -- type of handler
- **userfunc** (*function*) -- user function

This function defines the behavior of the library when an error occurs during the execution of the library's functions. This function should be (not mandatory) called before any other functions of the library. The behavior depends on the value of *typehandler*.

The possible values for *typehandler* are :

value	meaning
1	The library displays a message and continues the execution. The functions return an error code. The python and Octave/Matlab interfaces raise an exception. This is the default behavior of the library.
2	The library displays a message and terminates the execution with a system call to the function <i>exit</i> .
3	The library calls the user function <i>userfunc</i> with the message.

If the function is called with 1 or 2 for *typehandler*, the parameter *userfunc* must be set to 0.

The function *userfunc* must be defined as

```
def userfunc (msg)
# parameter msg is of type str
```

## MISCELLANEOUS FUNCTIONS

### 6.1 `calcephpy.getmaxsupportedorder`

`getmaxsupportedorder(segid)`

**Parameters**

**`segid`** (*int*) -- type of the segment.

**Returns**

version of the library

**Return type**

str

This function returns the maximal order of the derivatives computed by the functions `calcephpy.CalcephBin.compute_order()`, `calcephpy.CalcephBin.orient_order()`, .... for the segment type *segid*. If the segment type is unknown by the library, the function returns -1.

The accepted values of *segid*\* are the predefined constants `Constants.SEGTYPE_...` (*Constants*).

```
from calcephpy import *
maxorder = getmaxsupportedorder(Constants.SEGTYPE_SPK_2)
print('maximal order is ', maxorder)
```

### 6.2 `calcephpy.getversion_str`

`calcephpy.getversion_str()`

**Returns**

version of the library

**Return type**

str

This function returns the version of the CALCEPH Library, as a string.

```
from calcephpy import *
print('version=', getversion_str())
```



## NAIF IDENTIFICATION NUMBERS

The following predefined values must be used as the target body and origin of the coordinate system with the functions `calcephpy.CalcephBin.compute_unit()`, `calcephpy.CalcephBin.orient_unit()`, `calcephpy.CalcephBin.compute_order()` or `calcephpy.CalcephBin.orient_order()` if and only if the value `USE_NAIFID` has been set in the parameter `unit`.

This list is already predefined in the class `NaifId` (*Types*) of the module `calcephpy` (*Modules*) for the Python 2/3 interface. Relative to C or Fortran interface, the prefix `NAIFID_` is deleted for the following numbers.

### 7.1 Sun and planetary barycenters

Predefined Macros	NAIF ID	Name
NAIFID_SOLAR_SYSTEM_BARYCENTER	0	Solar System Barycenter
NAIFID_MERCURY_BARYCENTER	1	Mercury Barycenter
NAIFID_VENUS_BARYCENTER	2	Venus Barycenter
NAIFID_EARTH_MOON_BARYCENTER	3	Earth-Moon Barycenter
NAIFID_MARS_BARYCENTER	4	Mars Barycenter
NAIFID_JUPITER_BARYCENTER	5	Jupiter Barycenter
NAIFID_SATURN_BARYCENTER	6	Saturn Barycenter
NAIFID_URANUS_BARYCENTER	7	Uranus Barycenter
NAIFID_NEPTUNE_BARYCENTER	8	Neptune Barycenter
NAIFID_PLUTO_BARYCENTER	9	Pluto Barycenter
NAIFID_SUN	10	Sun

### 7.2 Coordinate Time ephemerides

Predefined Macros	NAIF ID	Name
NAIFID_TIME_CENTER	1000000000	center ID for Coordinate Time ephemerides <sup>1</sup>
NAIFID_TIME_TTMTDB	1000000001	Coordinate Time ephemeride TT-TDB <sup>2</sup>
NAIFID_TIME_TCGMTCB	1000000002	Coordinate Time ephemeride TCG-TCB <sup>2</sup>

---

<sup>1</sup> These values must only be used as a center body.

<sup>2</sup> These values must only be used as a target body.

## 7.3 Planet centers and satellites

Predefined Macros	NAIF ID	Name
NAIFID_MERCURY	199	Mercury
NAIFID_VENUS	299	Venus
NAIFID_EARTH	399	Earth
NAIFID_MOON	301	Moon
NAIFID_MARS	499	Mars
NAIFID_PHOBOS	401	Phobos
NAIFID_DEIMOS	402	Deimos
NAIFID_JUPITER	599	Jupiter
NAIFID_IO	501	Io
NAIFID_EUROPA	502	Europa
NAIFID_GANYMEDE	503	Ganymede
NAIFID_CALLISTO	504	Callisto
NAIFID_AMALTHEA	505	Amalthea
NAIFID_HIMALIA	506	Himalia
NAIFID_ELARA	507	Elara
NAIFID_PASIPHAЕ	508	Pasiphae
NAIFID_SINOPE	509	Sinope
NAIFID_LYSITHEA	510	Lysithea
NAIFID_CARME	511	Carme
NAIFID_ANANKE	512	Ananke
NAIFID_LEDA	513	Leda
NAIFID_THEBE	514	Thebe
NAIFID_ADRASTEА	515	Adrastea
NAIFID_METIS	516	Metis
NAIFID_CALLIRRHOЕ	517	Callirrhoe
NAIFID_THEMISTO	518	Themisto
NAIFID_MEGACLITE	519	Megaclite
NAIFID_TAYGETE	520	Taygete
NAIFID_CHALDENE	521	Chaldene
NAIFID_HARPALYKE	522	Harpalyke
NAIFID_KALYKE	523	Kalyke
NAIFID_IOCASTE	524	Iocaste
NAIFID_ERINOME	525	Erinome
NAIFID_ISONOE	526	Isonoe
NAIFID_PRAXIDIKE	527	Praxidike
NAIFID_AUTONOE	528	Autonoe
NAIFID_THYONE	529	Thyone
NAIFID_HERMIPPE	530	Hermippe
NAIFID_AITNE	531	Aitne
NAIFID_EURYDOME	532	Eurydome
NAIFID_EUANTHE	533	Euanthe
NAIFID_EUPORIE	534	Euporie
NAIFID_ORTHOSIE	535	Orthosie
NAIFID_SPONDE	536	Sponde

continues on next page

Table 1 – continued from previous page

Predefined Macros	NAIF ID	Name
NAIFID_KALE	537	Kale
NAIFID_PASITHEE	538	Pasithee
NAIFID_HEGEMONE	539	Hegemone
NAIFID_MNEME	540	Mneme
NAIFID_AOEDE	541	Aoede
NAIFID_THELXINOE	542	Thelxinoe
NAIFID_ARCHE	543	Arche
NAIFID_KALLICHORE	544	Kallichore
NAIFID_HELIKE	545	Helike
NAIFID_CARPO	546	Carpo
NAIFID_EUKELADE	547	Eukelade
NAIFID_CYLLENE	548	Cyllene
NAIFID_KORE	549	Kore
NAIFID_HERSE	550	Herse
NAIFID_DIA	553	Dia
NAIFID_SATURN	699	Saturn
NAIFID_MIMAS	601	Mimas
NAIFID_ENCELADUS	602	Enceladus
NAIFID_TETHYS	603	Tethys
NAIFID_DIONE	604	Dione
NAIFID_RHEA	605	Rhea
NAIFID_TITAN	606	Titan
NAIFID_HYPERION	607	Hyperion
NAIFID_IAPETUS	608	Iapetus
NAIFID_PHOEBE	609	Phoebe
NAIFID_JANUS	610	Janus
NAIFID_EPIMETHEUS	611	Epimetheus
NAIFID_HELENE	612	Helene
NAIFID_TELESTO	613	Telesto
NAIFID_CALYPSO	614	Calypso
NAIFID_ATLAS	615	Atlas
NAIFID_PROMETHEUS	616	Prometheus
NAIFID_PANDORA	617	Pandora
NAIFID_PAN	618	Pan
NAIFID_YMIR	619	Ymir
NAIFID_PAALIAQ	620	Paaliaq
NAIFID_TARVOS	621	Tarvos
NAIFID_IJIRAQ	622	Ijiraq
NAIFID_SUTTUNGR	623	Suttungr
NAIFID_KIVIUQ	624	Kiviuq
NAIFID_MUNDILFARI	625	Mundilfari
NAIFID_ALBIORIX	626	Albiorix
NAIFID_SKATHI	627	Skathi
NAIFID_ERRIAPUS	628	Erriapus
NAIFID_SIARNAQ	629	Siarnaq
NAIFID_THRYMR	630	Thrymr
NAIFID_NARVI	631	Narvi
NAIFID_METHONE	632	Methone
NAIFID_PALLENE	633	Pallene

continues on next page

Table 1 – continued from previous page

Predefined Macros	NAIF ID	Name
NAIFID_POLYDEUCES	634	Polydeuces
NAIFID_DAPHNIS	635	Daphnis
NAIFID_AEGIR	636	Aegir
NAIFID_BEBHIONN	637	Bebhionn
NAIFID_BERGELMIR	638	Bergelmir
NAIFID_BESTLA	639	Bestla
NAIFID_FARBAUTI	640	Farbauti
NAIFID_FENRIR	641	Fenrir
NAIFID_FORNJOT	642	Fornjot
NAIFID_HATI	643	Hati
NAIFID_HYROKKIN	644	Hyrokkin
NAIFID_KARI	645	Kari
NAIFID_LOGE	646	Loge
NAIFID_SKOLL	647	Skoll
NAIFID_SURTUR	648	Surtur
NAIFID_ANTHE	649	Anthe
NAIFID_JARNSAXA	650	Jarnsaxa
NAIFID_GREIP	651	Greip
NAIFID_TARQEQ	652	Tarqeq
NAIFID_AEGAEON	653	Aegaeon
NAIFID_URANUS	799	Uranus
NAIFID_ARIEL	701	Ariel
NAIFID_UMBRIEL	702	Umbriel
NAIFID_TITANIA	703	Titania
NAIFID_OBERON	704	Oberon
NAIFID_MIRANDA	705	Miranda
NAIFID_CORDELIA	706	Cordelia
NAIFID_OPHELIA	707	Ophelia
NAIFID_BIANCA	708	Bianca
NAIFID_CRESSIDA	709	Cressida
NAIFID_DESDEMONA	710	Desdemona
NAIFID_JULIET	711	Juliet
NAIFID_PORTIA	712	Portia
NAIFID_ROSALIND	713	Rosalind
NAIFID_BELINDA	714	Belinda
NAIFID_PUCK	715	Puck
NAIFID_CALIBAN	716	Caliban
NAIFID_SYCORAX	717	Sycorax
NAIFID_PROSPERO	718	Prospero
NAIFID_SETEBOS	719	Setebos
NAIFID_STEPHANO	720	Stephano
NAIFID_TRINCULO	721	Trinculo
NAIFID_FRANCISCO	722	Francisco
NAIFID_MARGARET	723	Margaret
NAIFID_FERDINAND	724	Ferdinand
NAIFID_PERDITA	725	Perdita
NAIFID_MAB	726	Mab
NAIFID_CUPID	727	Cupid

continues on next page

Table 1 – continued from previous page

Predefined Macros	NAIF ID	Name
NAIFID_NEPTUNE	899	Neptune
NAIFID_TRITON	801	Triton
NAIFID_NEREID	802	Nereid
NAIFID_NAIAD	803	Naiad
NAIFID_THALASSA	804	Thalassa
NAIFID_DESPINA	805	Despina
NAIFID_GALATEA	806	Galatea
NAIFID_LARISSA	807	Larissa
NAIFID_PROTEUS	808	Proteus
NAIFID_HALIMEDE	809	Halimede
NAIFID_PSAMATHE	810	Psamathe
NAIFID_SAO	811	Sao
NAIFID_LAOMEDEIA	812	Laomedeia
NAIFID_NESO	813	Neso
NAIFID_PLUTO	999	Pluto
NAIFID_CHARON	901	Charon
NAIFID_NIX	902	Nix
NAIFID_HYDRA	903	Hydra
NAIFID_KERBEROS	904	Kerberos
NAIFID_STYX	905	Styx

## 7.4 Comets

Predefined Macros	NAIF ID	Name
NAIFID_AREND	1000001	Arend
NAIFID_AREND_RIGAUX	1000002	Arend-Rigaux
NAIFID_ASHBROOK_JACKSON	1000003	Ashbrook-Jackson
NAIFID_BOETHIN	1000004	Boethin
NAIFID_BORRELLY	1000005	Borrelly
NAIFID_BOWELL_SKIFF	1000006	Bowell-Skiff
NAIFID_BRADFIELD	1000007	Bradfield
NAIFID_BROOKS_2	1000008	Brooks 2
NAIFID_BRORSEN_METCALF	1000009	Brorsen-Metcalf
NAIFID_BUS	1000010	Bus
NAIFID_CHERNYKH	1000011	Chernykh
NAIFID_CHURYUMOV_GERASIMENKO	1000012	Churyumov-Gerasimenko
NAIFID_CIFFREO	1000013	Ciffreo
NAIFID_CLARK	1000014	Clark
NAIFID_COMAS_SOLA	1000015	Comas Sola
NAIFID_CROMMELIN	1000016	Crommelin
NAIFID_D_ARREST	1000017	D"Drrest
NAIFID_DANIEL	1000018	Daniel
NAIFID_DE_VICO_SWIFT	1000019	De Vico-Swift
NAIFID_DENNING_FUJIKAWA	1000020	Denning-Fujikawa
NAIFID_DU_TOIT_1	1000021	Du Toit 1
NAIFID_DU_TOIT_HARTLEY	1000022	Du Toit-Hartley

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Table 2 – continued from previous page

Predefined Macros	NAIF ID	Name
NAIFID_DUTOIT_NEUJMIN_DELPORTE	1000023	Dutoit-Neujmin-Delporte
NAIFID_DUBIAGO	1000024	Dubiago
NAIFID_ENCKE	1000025	Encke
NAIFID_FAYE	1000026	Faye
NAIFID_FINLAY	1000027	Finlay
NAIFID_FORBES	1000028	Forbes
NAIFID_GEHRELS_1	1000029	Gehrels 1
NAIFID_GEHRELS_2	1000030	Gehrels 2
NAIFID_GEHRELS_3	1000031	Gehrels 3
NAIFID_GIACOBINI_ZINNER	1000032	Giacobini-Zinner
NAIFID_GICLAS	1000033	Giclas
NAIFID_GRIGG_SKJELLERUP	1000034	Grigg-Skjellerup
NAIFID_GUNN	1000035	Gunn
NAIFID_HALLEY	1000036	Halley
NAIFID_HANEDA_CAMPOS	1000037	Haneda-Campos
NAIFID_HARRINGTON	1000038	Harrington
NAIFID_HARRINGTON_ABELL	1000039	Harrington-Abell
NAIFID_HARTLEY_1	1000040	Hartley 1
NAIFID_HARTLEY_2	1000041	Hartley 2
NAIFID_HARTLEY_IRAS	1000042	Hartley-Iras
NAIFID_Herschel_RIGOLLET	1000043	Herschel-Rigollet
NAIFID_HOLMES	1000044	Holmes
NAIFID_HONDA_MRKOS_PAJDUSAKOVA	1000045	Honda-Mrkos-Pajdusakova
NAIFID_HOWELL	1000046	Howell
NAIFID_IRAS	1000047	Iras
NAIFID_JACKSON_NEUJMIN	1000048	Jackson-Neujmin
NAIFID_JOHNSON	1000049	Johnson
NAIFID_KEARNS_KWEE	1000050	Kearns-Kwee
NAIFID_KLEMOLA	1000051	Klemola
NAIFID_KOHOUTEK	1000052	Kohoutek
NAIFID_KOJIMA	1000053	Kojima
NAIFID_KOPFF	1000054	Kopff
NAIFID_KOWAL_1	1000055	Kowal 1
NAIFID_KOWAL_2	1000056	Kowal 2
NAIFID_KOWAL_MRKOS	1000057	Kowal-Mrkos
NAIFID_KOWAL_VAVROVA	1000058	Kowal-Vavrova
NAIFID_LONGMORE	1000059	Longmore
NAIFID_LOVAS_1	1000060	Lovas 1
NAIFID_MACHHOLZ	1000061	Machholz
NAIFID_MAURY	1000062	Maury
NAIFID_NEUJMIN_1	1000063	Neujmin 1
NAIFID_NEUJMIN_2	1000064	Neujmin 2
NAIFID_NEUJMIN_3	1000065	Neujmin 3
NAIFID_OLBERS	1000066	Olbers
NAIFID_PETERS_HARTLEY	1000067	Peters-Hartley
NAIFID_PONS_BROOKS	1000068	Pons-Brooks
NAIFID_PONS_WINNECKE	1000069	Pons-Winnecke
NAIFID_REINMUTH_1	1000070	Reinmuth 1
NAIFID_REINMUTH_2	1000071	Reinmuth 2
NAIFID_RUSSELL_1	1000072	Russell 1

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Predefined Macros	NAIF ID	Name
NAIFID_RUSSELL_2	1000073	Russell 2
NAIFID_RUSSELL_3	1000074	Russell 3
NAIFID_RUSSELL_4	1000075	Russell 4
NAIFID_SANGUIN	1000076	Sanguin
NAIFID_SCHAUMASSE	1000077	Schaumasse
NAIFID_SCHUSTER	1000078	Schuster
NAIFID_SCHWASSMANN_WACHMANN_1	1000079	Schwassmann-Wachmann 1
NAIFID_SCHWASSMANN_WACHMANN_2	1000080	Schwassmann-Wachmann 2
NAIFID_SCHWASSMANN_WACHMANN_3	1000081	Schwassmann-Wachmann 3
NAIFID_SHAJN_SCHALDACH	1000082	Shajn-Schaldach
NAIFID_SHOEMAKER_1	1000083	Shoemaker 1
NAIFID_SHOEMAKER_2	1000084	Shoemaker 2
NAIFID_SHOEMAKER_3	1000085	Shoemaker 3
NAIFID_SINGER_BREWSTER	1000086	Singer-Brewster
NAIFID_SLAUGHTER_BURNHAM	1000087	Slaughter-Burnham
NAIFID_SMIRNOVA_CHERNYKH	1000088	Smirnova-Chernykh
NAIFID_STEPHAN_OTERMA	1000089	Stephan-Oterma
NAIFID_SWIFT_GEHRELS	1000090	Swift-Gehrels
NAIFID_TAKAMIZAWA	1000091	Takamizawa
NAIFID_TAYLOR	1000092	Taylor
NAIFID_TEMPEL_1	1000093	Tempel 1
NAIFID_TEMPEL_2	1000094	Tempel 2
NAIFID_TEMPEL_TUTTLE	1000095	Tempel-Tuttle
NAIFID_TRITTON	1000096	Tritton
NAIFID_TSUCHINSHAN_1	1000097	Tsuchinshan 1
NAIFID_TSUCHINSHAN_2	1000098	Tsuchinshan 2
NAIFID_TUTTLE	1000099	Tuttle
NAIFID_TUTTLE_GIACOBINI_KRESAK	1000100	Tuttle-Giacobini-Kresak
NAIFID_VAISALA_1	1000101	Vaisala 1
NAIFID_VAN_BIESBROECK	1000102	Van Biesbroeck
NAIFID_VAN_HOUTEN	1000103	Van Houten
NAIFID_WEST_KOHOUTEK_IKEMURA	1000104	West-Kohoutek-Ikemura
NAIFID_WHIPPLE	1000105	Whipple
NAIFID_WILD_1	1000106	Wild 1
NAIFID_WILD_2	1000107	Wild 2
NAIFID_WILD_3	1000108	Wild 3
NAIFID_WIRTANEN	1000109	Wirtanen
NAIFID_WOLF	1000110	Wolf
NAIFID_WOLF_HARRINGTON	1000111	Wolf-Harrington
NAIFID_LOVAS_2	1000112	Lovas 2
NAIFID_URATA_NIIJIMA	1000113	Urata-Niijima
NAIFID_WISEMAN_SKIFF	1000114	Wiseman-Skiff
NAIFID_HELIN	1000115	Helin
NAIFID_MUELLER	1000116	Mueller
NAIFID_SHOEMAKER_HOLT_1	1000117	Shoemaker-Holt 1
NAIFID_HELIN_ROMAN_CROCKETT	1000118	Helin-Roman-Crockett
NAIFID_HARTLEY_3	1000119	Hartley 3
NAIFID_PARKER_HARTLEY	1000120	Parker-Hartley
NAIFID_HELIN_ROMAN_ALU_1	1000121	Helin-Roman-Alu 1
NAIFID_WILD_4	1000122	Wild 4

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Predefined Macros	NAIF ID	Name
NAIFID_MUELLER_2	1000123	Mueller 2
NAIFID_MUELLER_3	1000124	Mueller 3
NAIFID_SHOEMAKER_LEVY_1	1000125	Shoemaker-Levy 1
NAIFID_SHOEMAKER_LEVY_2	1000126	Shoemaker-Levy 2
NAIFID_HOLT_OLMSTEAD	1000127	Holt-Olmstead
NAIFID_METCALF_BREWINGTON	1000128	Metcalf-Brewington
NAIFID_LEVY	1000129	Levy
NAIFID_SHOEMAKER_LEVY_9	1000130	Shoemaker-Levy 9
NAIFID_HYAKUTAKE	1000131	Hyakutake
NAIFID_HALE_BOPP	1000132	Hale-Bopp
NAIFID_SIDING_SPRING	1003228	Siding Spring

## RELEASE NOTES

- **Version 3.5.2**

Fix a possible buffer overflow in the python interface of getfileversion.  
Fix the undefined external symbol rint with visual C compiler 10.0.  
Fix the redefinition of vasprintf with the environment MinGW.

- **Version 3.5.1**

Support the incremental assignment in the text SPICE kernel file (files .tpc).  
Support the relative frames in the frame SPICE kernel file (files .tf).  
Fix the returned values of calceph\_getconstantvs with large SPICE kernels.  
Fix random crash or double-free error when meta-kernel SPICE kernel files are closed (files .mk).  
Raise an error inside the function calceph\_compute if center != 0 for the target (14,15,16,17), to conform to the documentation.  
Fix the object name Megaclite (519) from NAIFID\_MAGACLITE to NAIFID\_MEGACLITE.

- **Version 3.5.0**

Add the function calceph\_getmaxsupportedorder, calceph\_getpositionrecordindex2 and calceph\_getorientrecordindex2.  
Update config.sub and config.guess to support new processors.  
Fix an incorrect result of calceph\_(s)getconstant and calceph\_getconstantvd with SPICE text kernel if the current locale has a decimal separator different than the decimal point (e.g. comma).  
Fix compilation warnings with gcc 11 about pointers and arrays.  
Fix the installation with python 3.10 or later (PEP 632).  
Fix compilation errors with the android SDK API level < 24.  
Reduce the stack size of the arrays for the evaluation with C99 compilers.  
Change the array size of the function calceph\_getfileversion from the value CALCEPH\_MAX\_CONSTANTNAME to CALCEPH\_MAX\_CONSTANTVALUE, to conform to the documentation and examples.  
Only C and Fortran 2003 are affected by this change : previous header file contains an invalid declaration of the function.  
Few programs may be affected by this change, as the required array has a smaller size : No buffer overflow may appear.

- **Version 3.4.7**

Fix a decode error of the little-endian SPICE kernel files on the big-endian architectures (e.g. processor s390x).

Fix the transmission of the flags FCFLAGS to the fortran compilers.

- **Version 3.4.6**

Fix a wrong error message about unsupported order for the segment 21.

Fix incorrect results for SPICE kernel files containing segments of type 21 with many records ( $\geq 100$ ) and improved the accuracy if segments of type 21 contain few records ( $< 100$ ).

- **Version 3.4.5**

Fix a random crash of `calceph_open_array` if one of the file is invalid.

`f90calceph_seterrorhandler` now ignores the parameter `userfunc`, instead of the requirement to set to 0, if the parameter type is 1 or 2. `userfunc` can be an empty function. It fixes compilation errors with gcc 10.1.

- **Version 3.4.4**

Fix a regression introduced in 3.4.3 (remove a recursion with SPICE kernel files).

- **Version 3.4.3**

Remove a recursion to read the segments of the SPICE kernel files. It reduces the usage of the stack.

Fix the installation of python package under Anaconda.

- **Version 3.4.2**

Add a missing makefile for windows system using the Visual C++ compiler.

Support SPICE kernels larger than 4GBytes.

- **Version 3.4.1**

Improve the execution time of `calceph_open` and `calceph_open_array` if the spice kernels contains a large number of bodies.

Update `config.sub` and `config.guess` to support arm processors.

- **Version 3.4.0**

Add the function `calceph_isthreadsafe`.

Multiple threads can now access the same ephemeris descriptor if the function `calceph_isthreadsafe` returns 1.

Fortran and C examples (`f2003parallel.f`, `cparallel.c`), written using OpenMP, are available in the folder examples.

Fix an error if multiple SPICE kernels are loaded for the same objects over different time-span.

Fix the MinGW Makefiles if the variable `MAKE` contains spaces.

Support the segment 5 and 18 in the SPICE kernel file.

Support the euler angles for the orientation stored in a text PCK files (`BODY..._POLE_RA`, `BODY..._POLE_DE`, `BODY..._POLE_PM`, `BODY..._NUT_PREC...`).

Support the frame 17 (ECLIPJ2000) in the SPICE kernel file.

Add the utilities `calceph_queryposition` and `calceph_queryorientation`.

- **Version 3.3.1**

Fix the installation with python 3.7.0 or later.  
 Fix the installation with python and pip on Windows operating system.  
 Add the missing file pythonapi/src/Makefile.mingw for the environment MinGW.

- **Version 3.3.0**

Add the functions `calceph_getfileversion`.  
 Fix a regression to open some old JPL DE format files.  
 Fix a compiler warning in the file `util.c`.  
 Support the segments 8, 9, 17 and 21 in the SPICE kernel file.  
 Check the validity of the number of constants in the original INPOP/DE files.  
 For the Python interface, the functions `compute???` and `orient???` supports now a list or numpy's array for the time parameters.

- **Version 3.2.0**

Fix the creation of the dynamic library with `msys/mingw` on Windows.  
 Fix the returned value of the functions `f90calceph_getconstantvd` and `f90calceph_getconstantvs`.  
 Fix a compilation warning with the GNU C compilers 8.0 or later.  
 Support the original JPL files with TT-TDB or with a large number of constants.  
 Support the IAU 1980 Nutation Angles of the JPL files.  
 Add the NAIF identification numbers for DIA, KERBEROS, STYX and SIDING SPRING.  
 Add the option `installnodoc` to the make command.

- **Version 3.1.0**

Add the Mex interface compliant with Octave 4.0+ and Matlab 2017+.  
 Add the functions `calceph_getconstantsd`, `calceph_getconstantvd` and `calceph_getconstantss` and `calceph_getconstantvs`.  
 Fix a compilation problem with MinGW if the terminal `cmd.exe` is used.  
 Fix a wrong function name `open_array` instead of `open` in the documentation of the Python interface.  
 Fix the return value of the functions `calceph_orient_xxx` when the unit `CALCEPH_UNIT_RAD` is not provided.  
 The return value of the function `calceph_(s)getconstant(index)` is the number of values associated to the constant.  
 Display a better message for the unsupported old spice kernel (NAIF/DAF)

- **Version 3.0.0**

Update the license CeCILL v2.0 to CeCILL v2.1.  
 Fix a decode error for SPICE kernels with a big-endian format.  
 Add the function `calceph_gettimescale` and `calceph_gettimespan`.  
 Add the function `calceph_getpositionrecordcount` and `calceph_getpositionrecordindex`.  
 Add the function `calceph_getorientrecordcount` and `calceph_getorientrecordindex`.  
 Add the function `calceph_sgettimescale` and `calceph_sgettimespan`.  
 Support INPOP file format 3.0 (add angular momentum due to the rotation in the binary file).  
 Use sphinx-doc to produce the documentation.

- **Version 2.3.2**

Fix the return value of the function `calceph_getconstant` if the constant name "AU" or "EMRAT" is not available.

Fix the documentation for the fortran interface of the function `calceph_prefetch`.

Fix the return value of the function `calceph_orient_unit` if the frame SPICE kernel file is missing.

- **Version 2.3.1**

Fix the compilation warnings with the Pelles compiler.

Fix the compilation warnings with the C89 standard.

Fix the compilation warnings with the GNU C compilers.

Fix the documentation for the constant `CALCEPH_VERSION_STRING`.

- **Version 2.3.0**

Add the python interface compliant with python 2.6+ and python 3.

Add the preprocessor macro `CALCEPH_VERSION_STRING`.

Add the function `calceph_getversion_str`.

Add the function `calceph_compute_order` and `calceph_orient_order`.

Fix the return value of the functions `calceph_compute_xxx` when the reference frame is not available in the spice kernel files.

The function should produce an error and return 0 (before the function performed no computation but it returned 1).

- **Version 2.2.5**

Fix an incorrect result if `CALCEPH_UNIT_DAY` is provided to `calceph_compute_unit` and the target is TCG-TCB or TT-TDB.

Support the numerical constants declared without parenthesis in the text kernel files (.tpc).

Support the segment 1, 12 and 13 in the SPICE kernel file.

- **Version 2.2.4**

Update the version number of the dynamic library.

- **Version 2.2.3**

Add the predefined constants for calceph version in the fortran interface.

Fix the build chain if calceph is compiled from another folder.

- **Version 2.2.2**

Support the compilation in the standard C89.

- **Version 2.2.1**

Remove debug informations that are printed when errors occur in `calceph_?compute_???`.

Support the Portland compilers.

Fix the info documentation.

Report an error if no asteroid is available in an ephemeris file with the INPOP file format (instead of a crash).

- **Version 2.2.0**

Support the new segments 20, 102, 103 and 120 in the SPICE kernel file.

Support the NAIF identification numbers.

Add the functions `calceph_orient_unit` and `calceph_prefetch`.

- **Version 2.1.0**

Fix a bug in `calceph_getconstant` and `calceph_sgetconstant` with an invalid name

Remove the null character in the name of the constant returned by the function

`(f90)calceph_(s)getconstantindex` when the Fortran interface is used.

- **Version 2.0.0**

Fix memory leaks in `calceph_open` when errors occur.

Support INPOP file format 2.0 (supports TCB ephemeris file and add asteroids in the binary file).

Add the function `calceph_open_array` and `calceph_compute_unit`.

Add the tools `calceph_inspector` to show details about ephemeris file.

Support SPICE kernel file (SPK with segment 2 or 3, text and binary PCK, meta kernel, basic frame kernel).

Improve the performances.

Correct the Fortran 2003 interface for `calceph_sgetconstantindex`.

Add the constant 17 to get TCG-TCB from TCB ephemeris file.

- **Version 1.2.0**

Change the licensing : triple licenses to support integration in BSD software.

Remove explicit dependencies on the record size for `DExxx`.

- **Version 1.1.2**

Fix a compilation warning with oracle studio compiler 12.

Fix a bug with gcc on solaris in 64 bit mode.

Fix the copyright statements.

- **Version 1.1.1**

Fix a compilation error in `util.h` and a warning with the sun studio compilers.

- **Version 1.1.0**

Add the function `calceph_seterrorhandler` for the custom error handlers.

- **Version 1.0.3**

Support the JPL ephemeris file DE423.

- **Version 1.0.2**

Fix memory leaks in the fortran-90 interface.

- **Version 1.0.1**

Support the large ephemeris files (>2GB) on 32-bit operating systems.

Fix the documentation of the function `f90calceph_sopen`.

Fix an invalid open mode on Windows operating systems.

Report accurately the I/O errors.

- **Version 1.0.0**

Initial release.

## REPORTING BUGS

If you think you have found a bug in the CALCEPH Library, first have a look on the CALCEPH Library web page <http://www.imcce.fr/inpop>, in which case you may find there a workaround for it. Otherwise, please investigate and report it. We have made this library available to you, and it seems very important for us, to ask you to report the bugs that you find.

There are a few things you should think about when you put your bug report together. You have to send us a test case that makes it possible for us to reproduce the bug. Include instructions on the way to run the test case.

You also have to explain what is wrong; if you get a crash, or if the results printed are incorrect and in that case, in what way.

Please include compiler version information in your bug report. This can be extracted using `cc -V` on some machines, or, if you're using gcc, `gcc -v`. Also, include the output from `uname -a` and the CALCEPH version.

Send your bug report to: [inpop.imcce@obspm.fr](mailto:inpop.imcce@obspm.fr). If you think something in this manual is unclear, or downright incorrect, or if the language needs to be improved, please send a note to the same address.



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